



EVLA Tests Mtg

Moon Compression Test Results

Rob Selina

Rick Perley



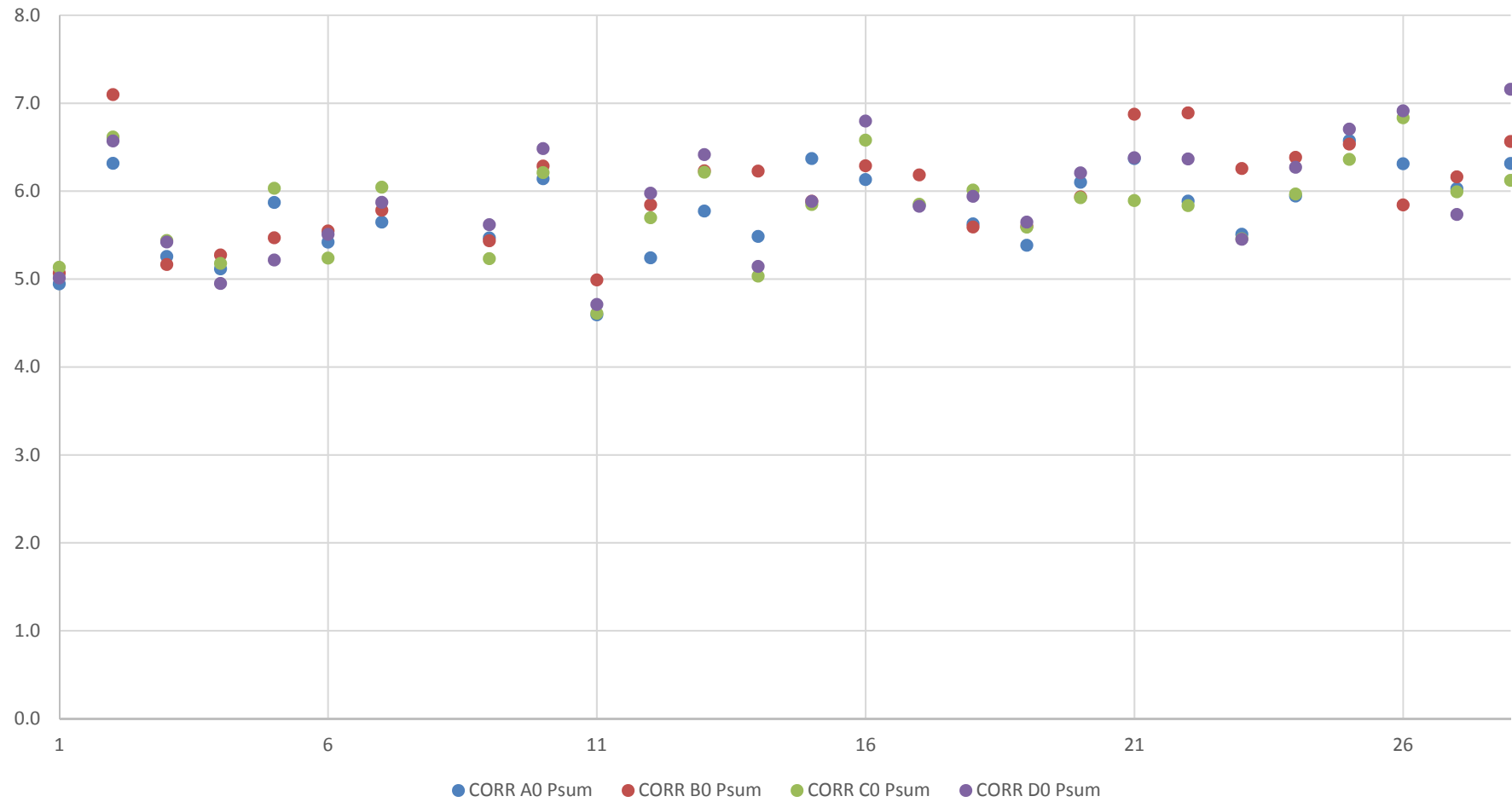
Test Goals

- Gain some insight into the source of the gain compression.
- Determine if the problem is primarily ‘analog’ or ‘digital’ in nature.
- Pick an intermediate point in the signal path and attempt to measure changes in switched power (P_{diff} , P_{sum}) at that point.
 - If results are consistent with the correlator derived switched power, problem is likely upstream of that that point.
- Use Total Power Detectors (TPDs) in T304. Secondary goal of finding new applications for these tools.
- Data collected on 04-07-2016

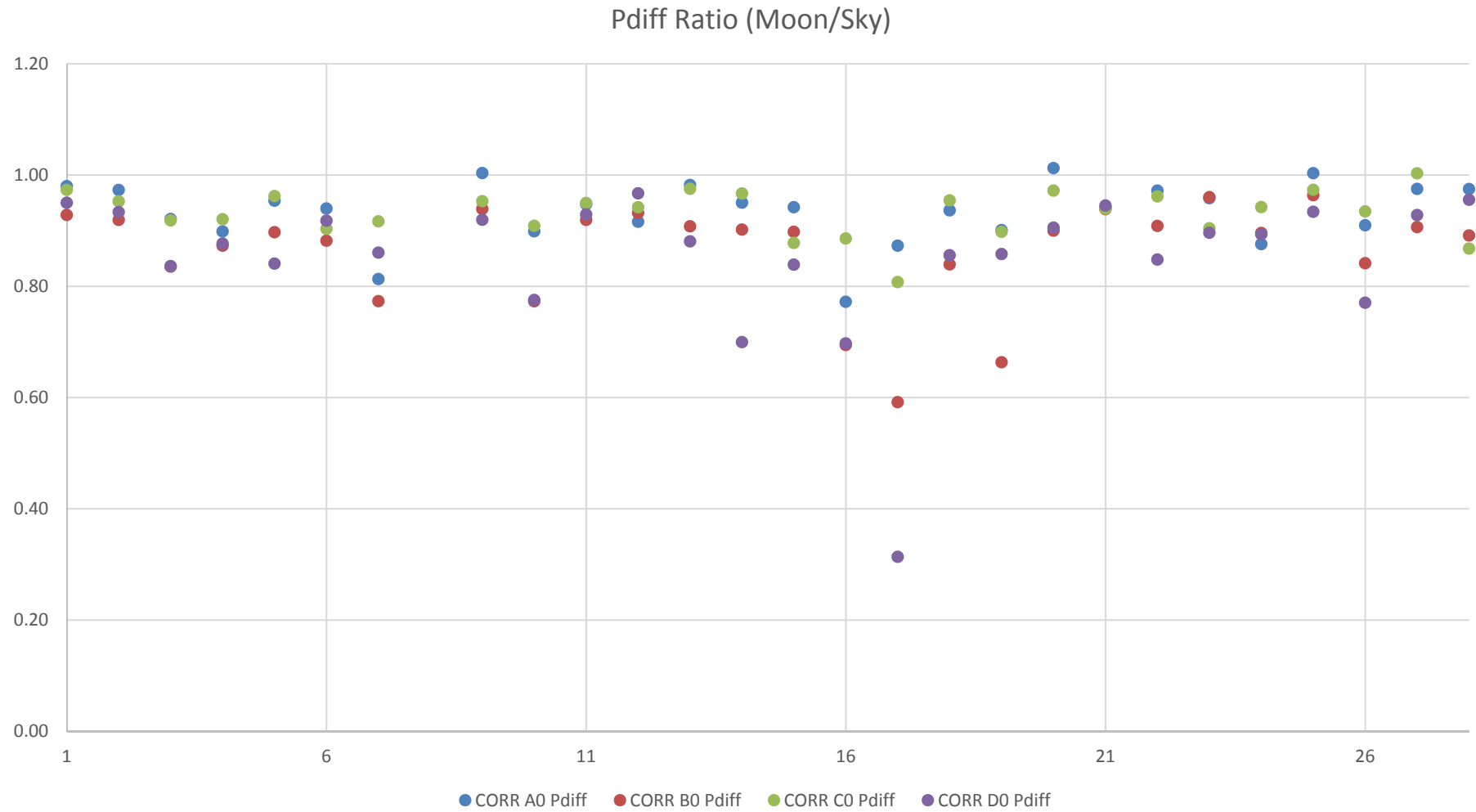


Array Performance – Gain Variations

Correlator Pmoon/Psky



Array Performance – Gain Compression

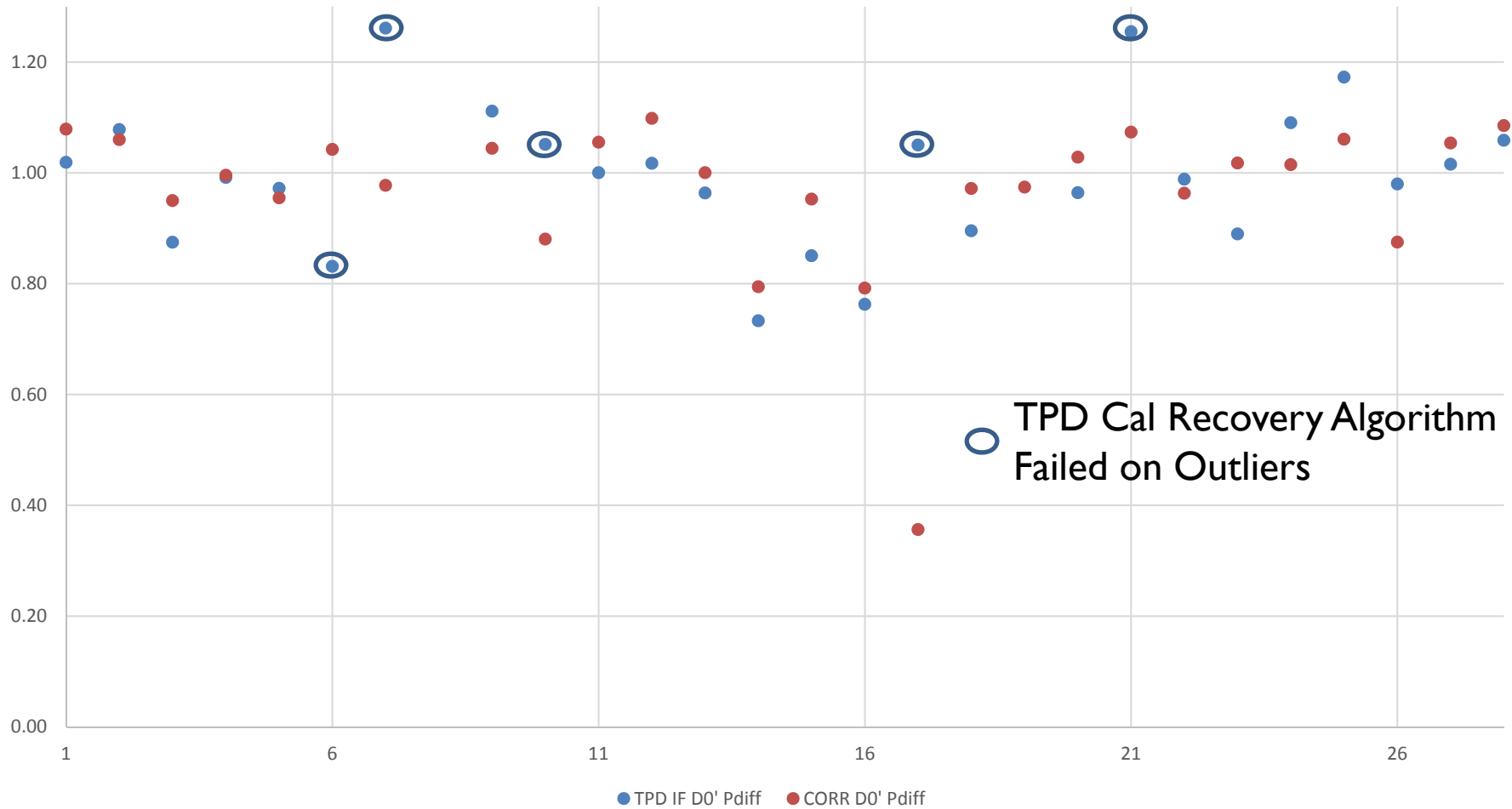


Gain Compression – TPDs vs. Correlator

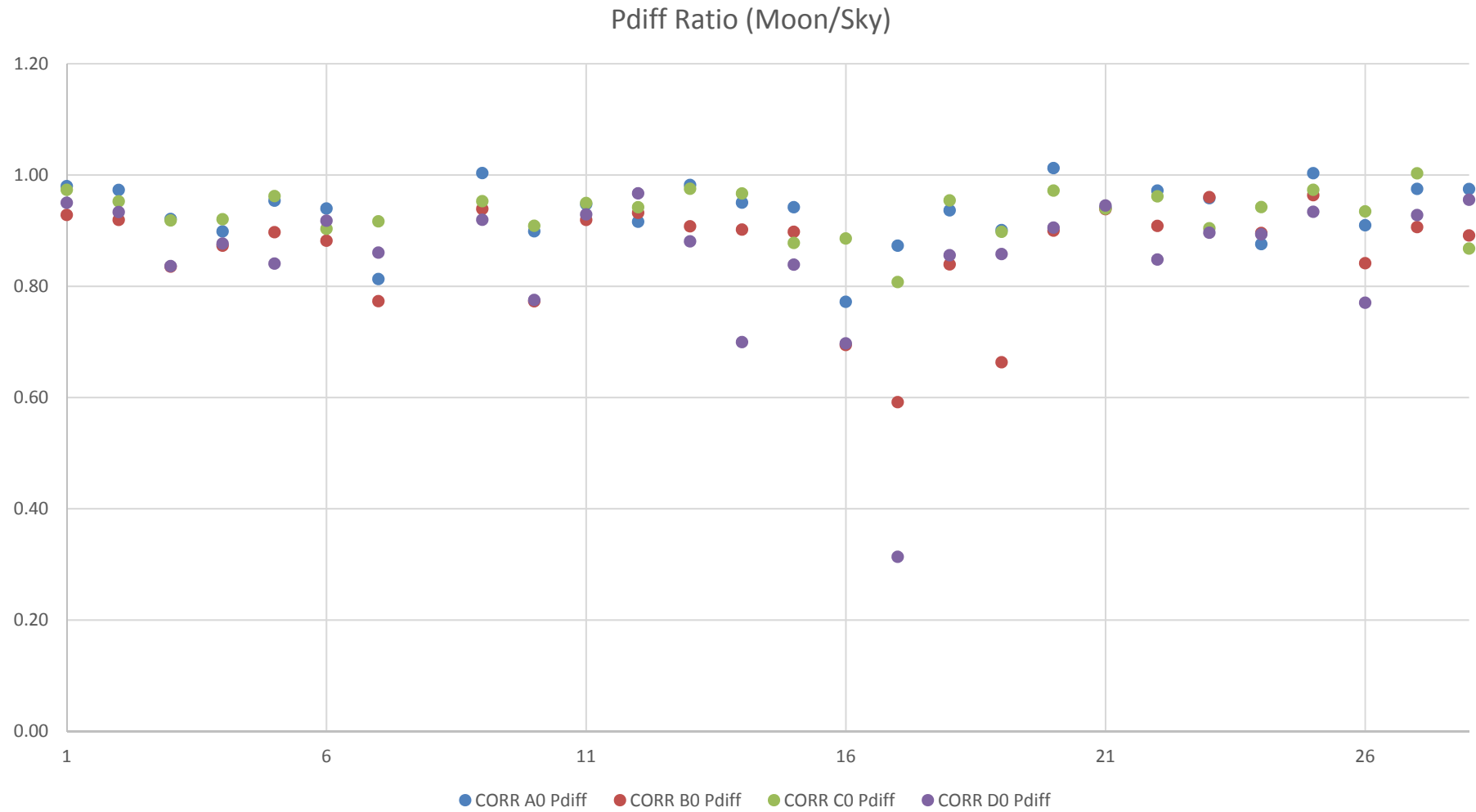
IF D

Correlator VS TPD

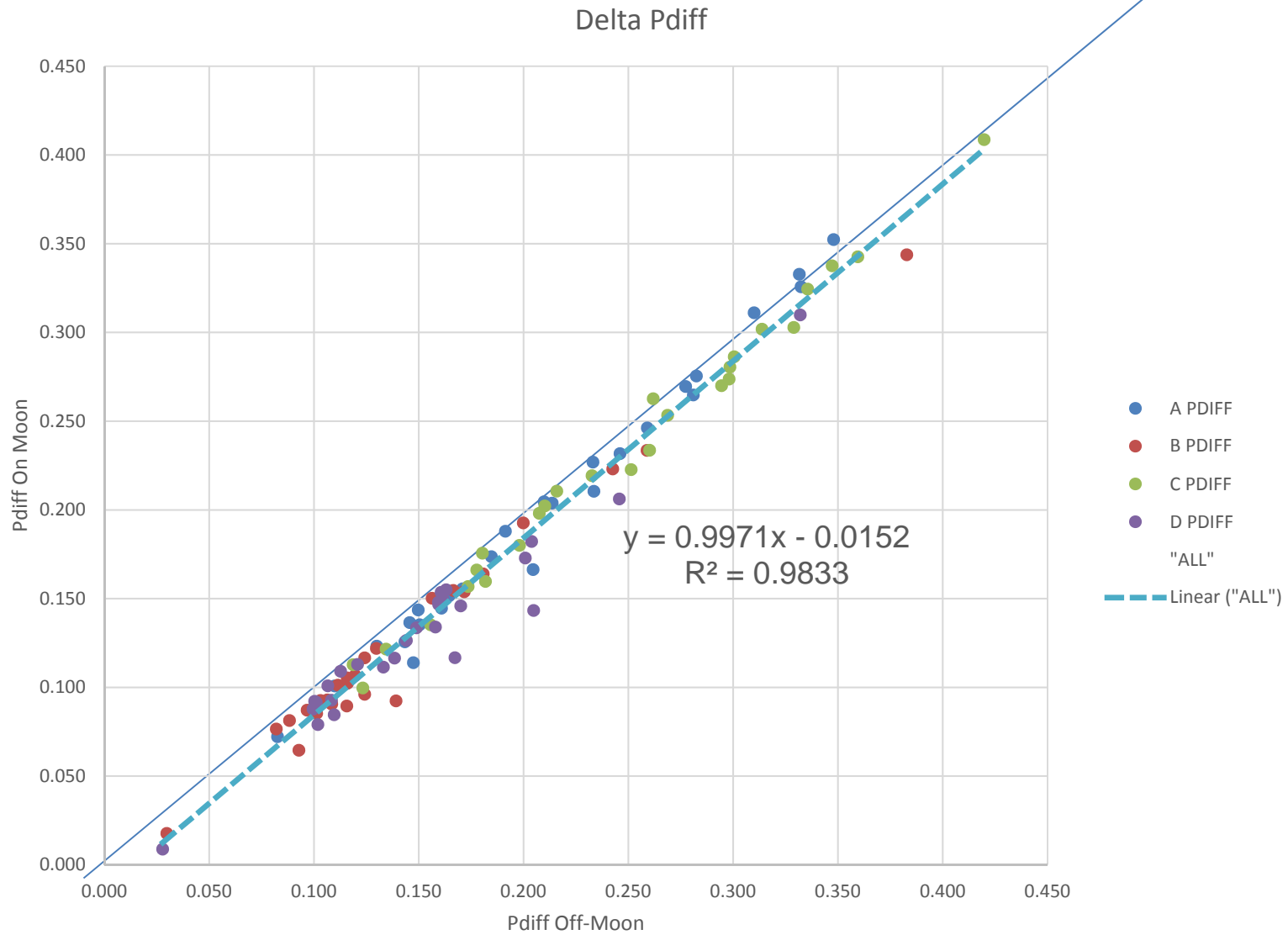
Pdiff Ratio Pmoon/Psky, **Normalized**



Array Performance – Gain Compression

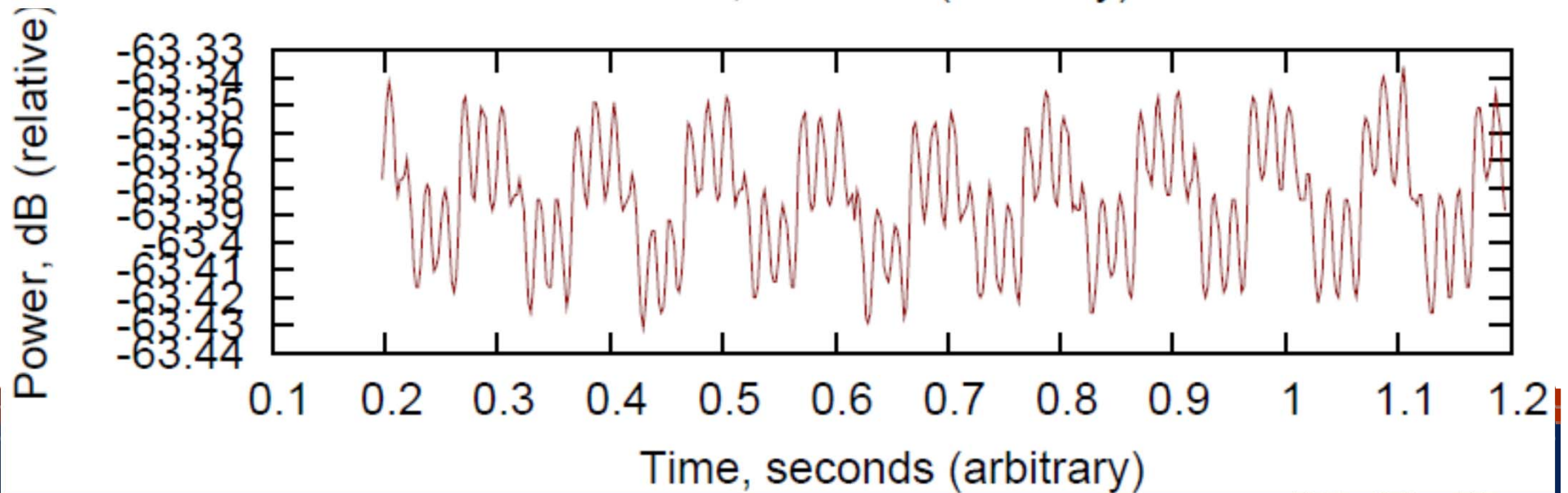
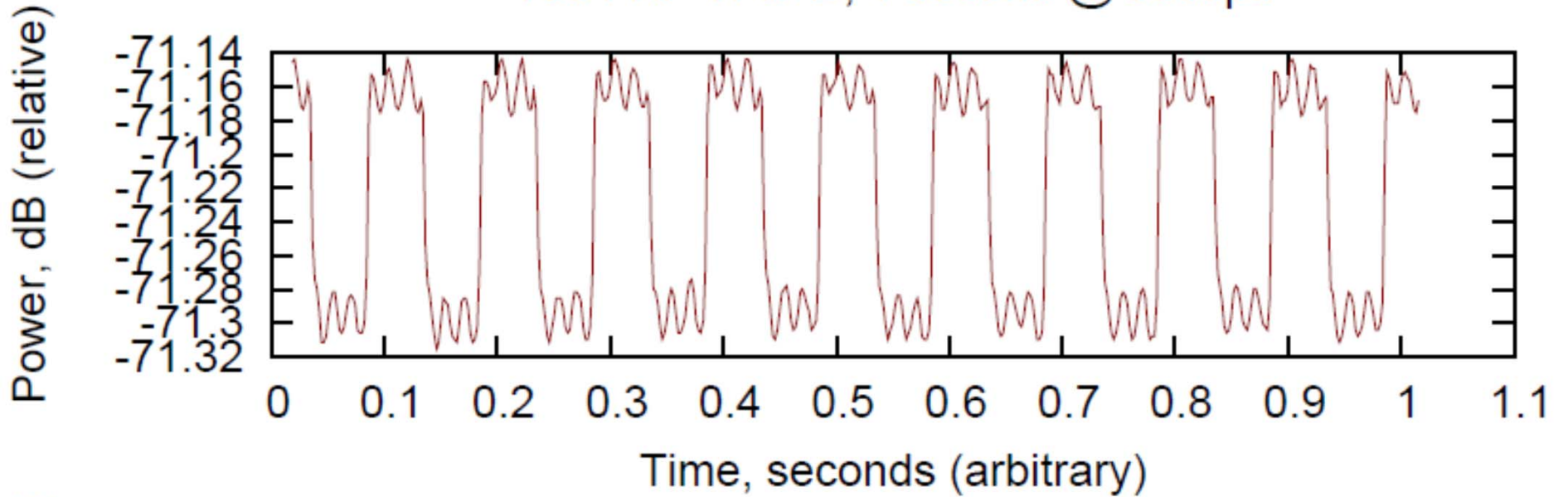


Delta Pdiff



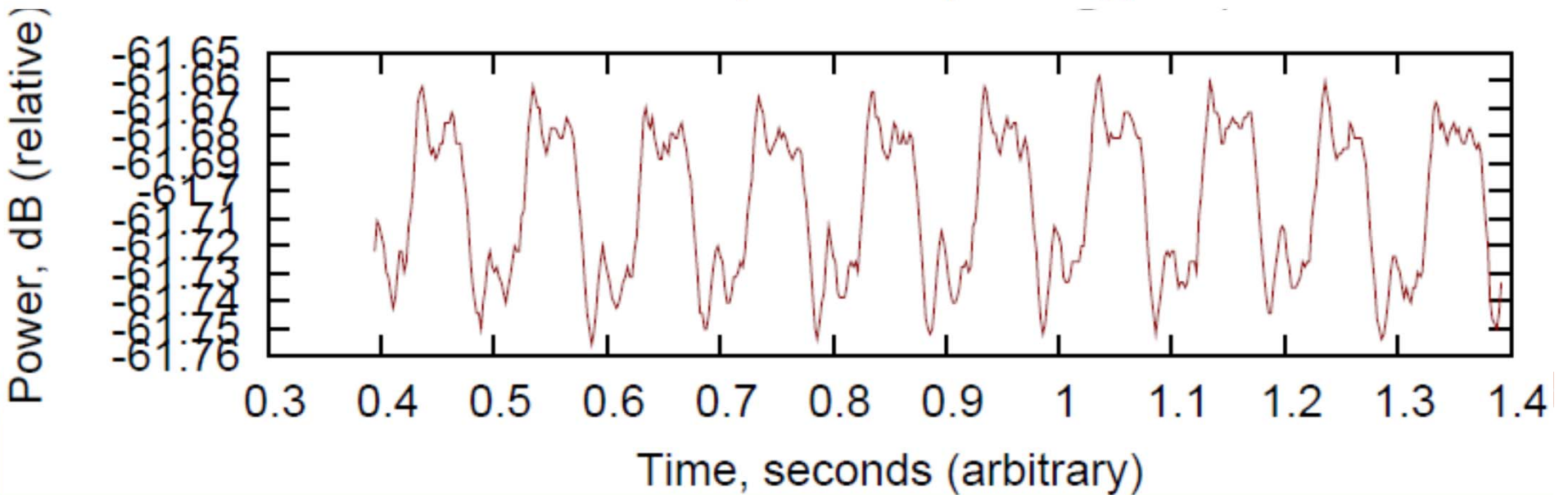
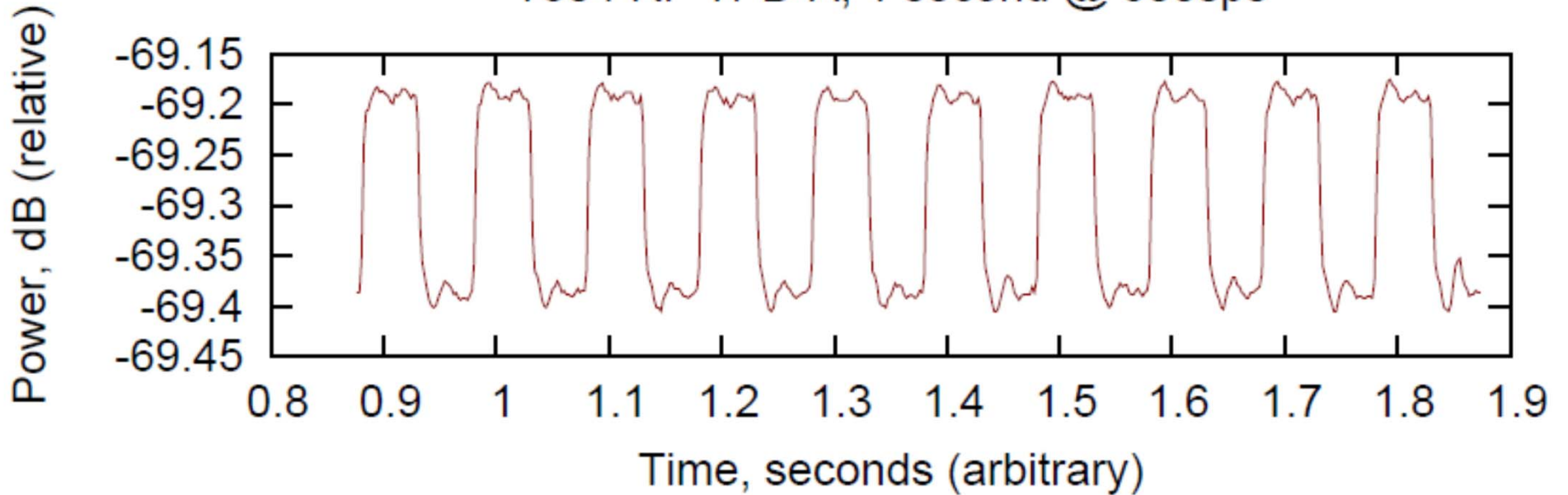
Gain Modulation – OFF/ON MOON (e.g EA06)

T304 RF TPD-D, 1 second @ 500sps



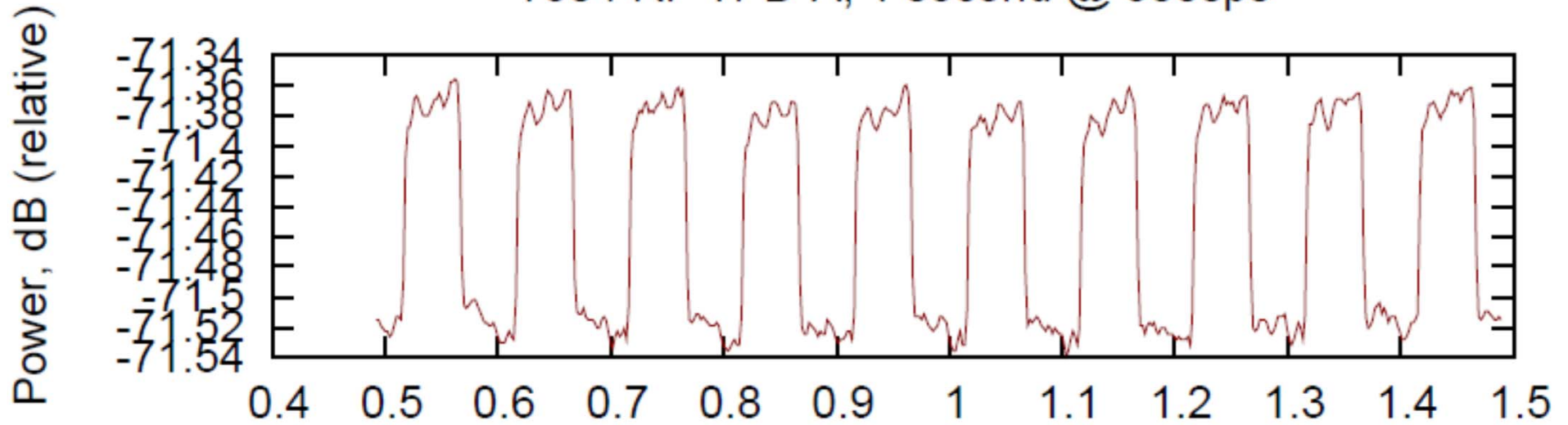
Gain Modulation – CAL Ringing (e.g EA02)

T304 RF TPD-A, 1 second @ 500sps

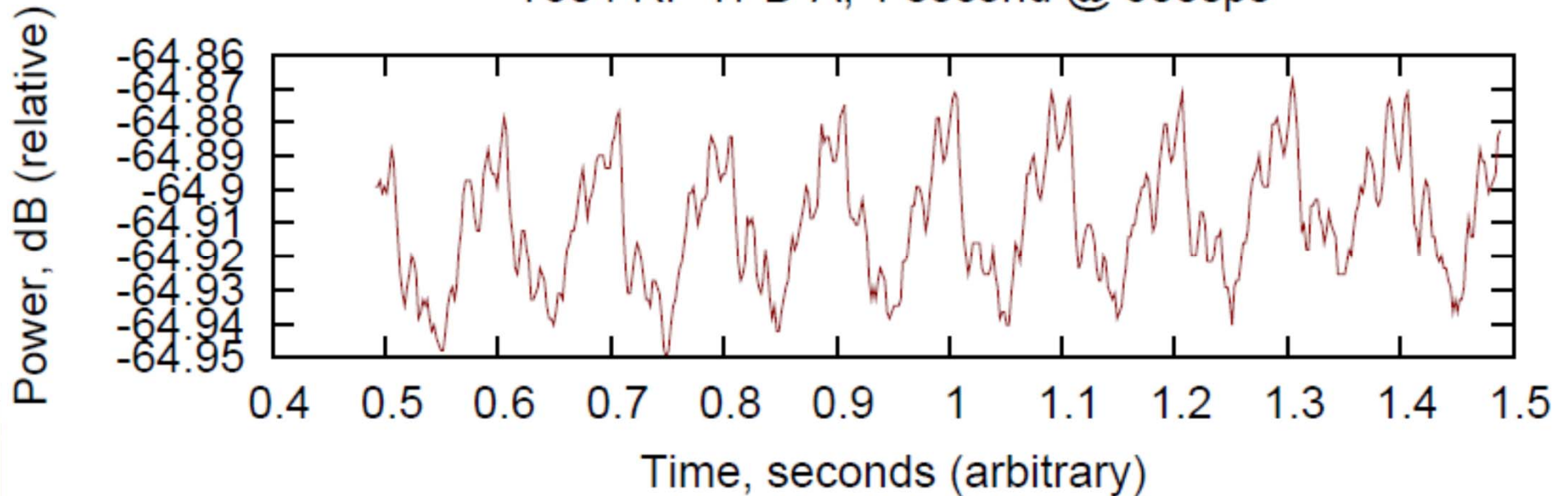


Gain Modulation (e.g EA07)

T304 RF TPD-A, 1 second @ 500sps



T304 RF TPD-A, 1 second @ 500sps



Pdiff with Gain Modulation

$$P_{off} = G_a (P_{ant} + P_{rx})$$

$$P_{on} = G_a (P_{ant} + P_{rx} + P_{cal})$$

$$P_{diff} = P_{on} - P_{off}$$

$$P_{diff} = G_a P_{cal}$$

...but P_{ant} seems to create more gain modulation, so P_{cal} may too:

$$P_{off} = G_{off} (P_{ant} + P_{rx})$$

$$P_{on} = G_{on} (P_{ant} + P_{rx} + P_{cal})$$

$$P_{diff} = G_{on} P_{cal} + (G_{off} - G_{on})(P_{ant} + P_{rx})$$

If $G \approx C + G_{mod}$, then $G_{off} - G_{on}$ is dominated by the change in gain modulation.

Pdiff with Gain Modulation

$$P_{diff} = G_{on} P_{cal} + (G_{off} - G_{on})(P_{ant} + P_{rx})$$

...if $P_{cal} \approx 3\% P_{sky}$, and $P_{moon} \approx 6P_{sky}$

$$G_{sky} \approx C + \langle G_{mod}(t) \rangle \quad G_{sky-off} \approx C + \langle G_{mod}(t) \rangle$$

$$G_{moon} \approx C + \langle 6G_{mod}(t) \rangle \quad G_{sky-on} \approx C + \langle 1.03G_{mod}(t) \rangle$$

$$G_{moon-off} \approx C + \langle 6G_{mod}(t) \rangle$$

$$G_{moon-on} \approx C + \langle 6.03G_{mod}(t) \rangle$$

$$P_{diff-sky} = C \cdot P_{cal} + \langle 1.03G_{mod}(t) \rangle P_{cal} + \langle 0.03G_{mod}(t) \rangle (P_{sky} + P_{rx})$$

$$P_{diff-moon} = C \cdot P_{cal} + \langle 6.03G_{mod}(t) \rangle P_{cal} + \langle 0.03G_{mod}(t) \rangle (6P_{sky} + P_{rx})$$



Delta Pdiff Model

- Real Data fit:

$$y = 0.9971x - 0.0152$$

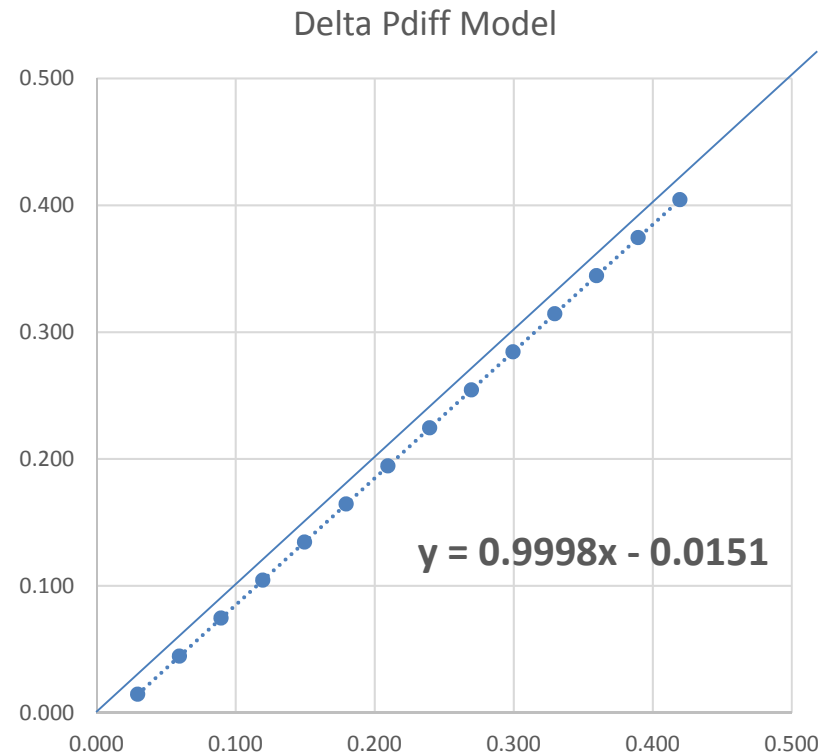
Let $\langle G_{\text{mod}} \rangle$ be:

-0.0018% for Cal ON

0.0018% for Cal OFF

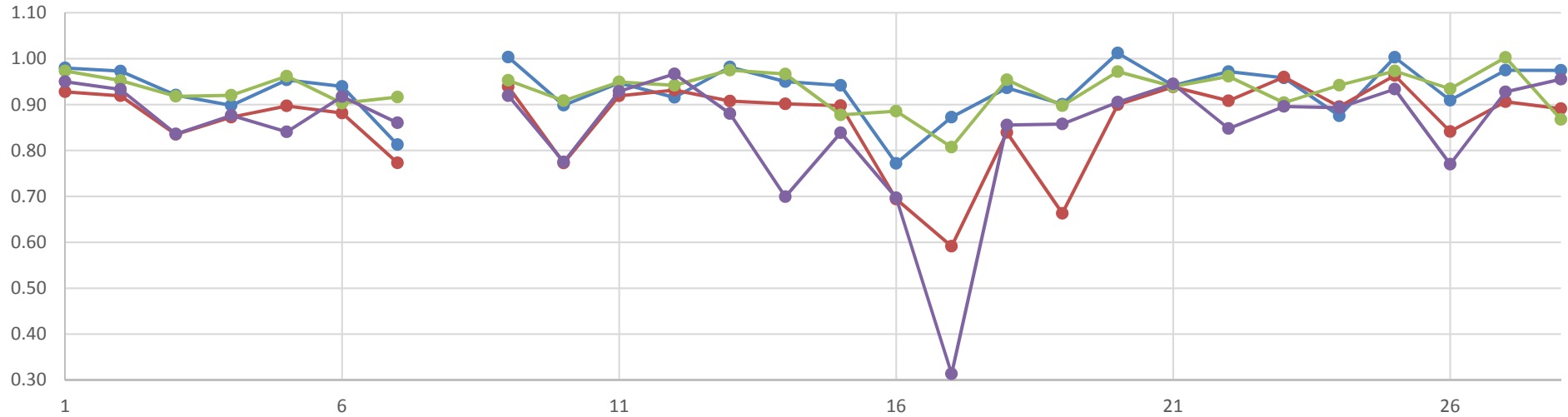
model Pdiff response:

$$y = 0.9998x - 0.0151$$

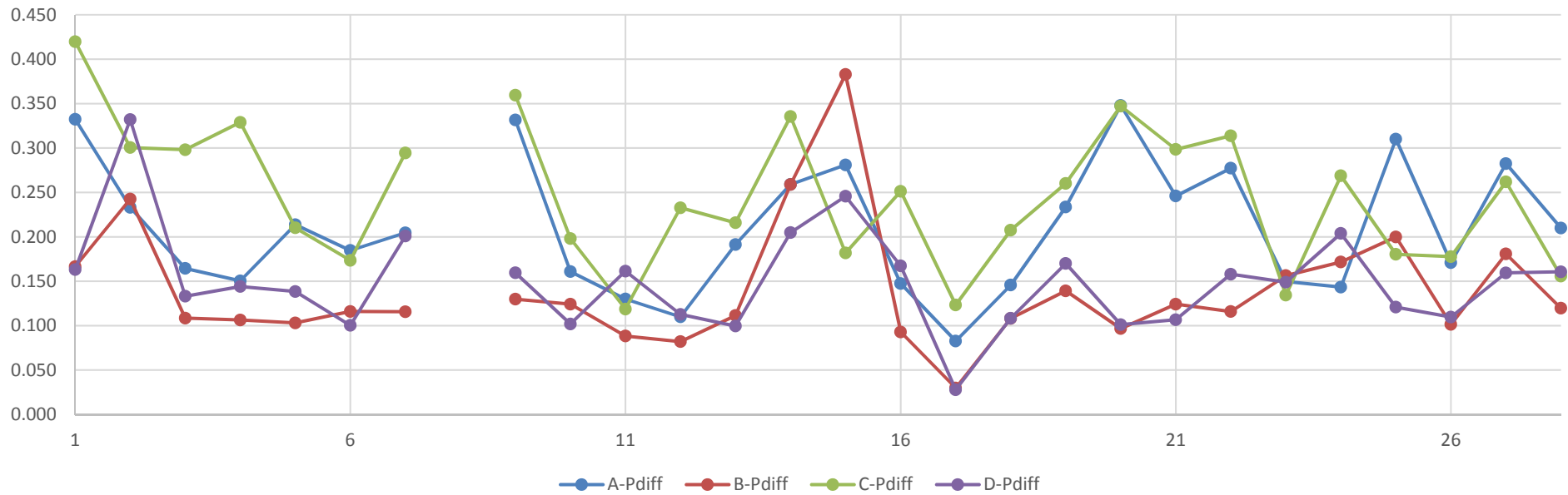


Gain Mod + Low Pcal = More Pdiff Compression

Pdiff Ratio (Moon/Sky)

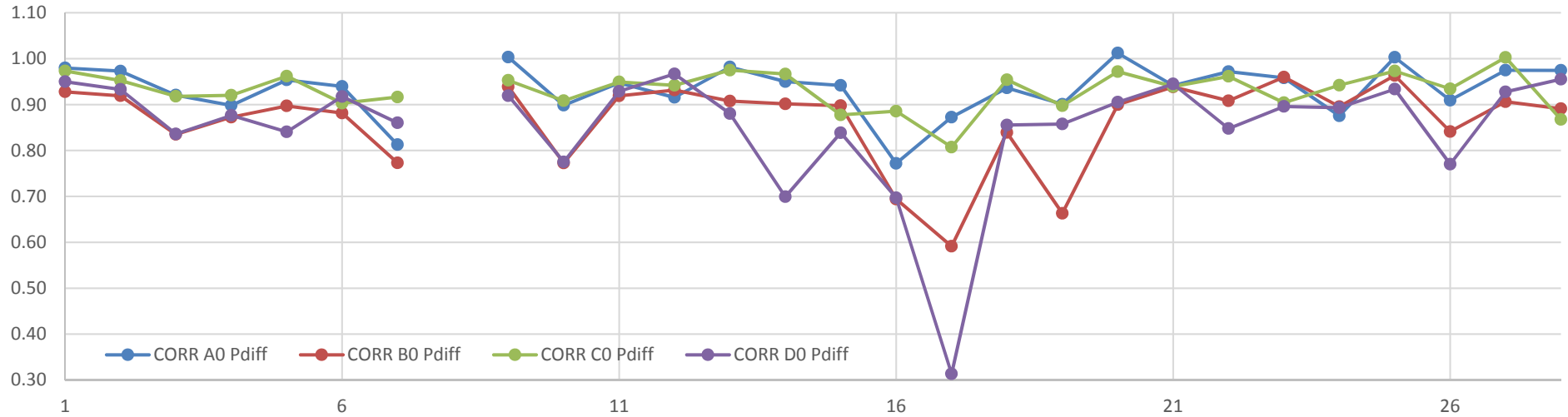


Cal Amplitude (Pdiff Off-Moon, Counts)

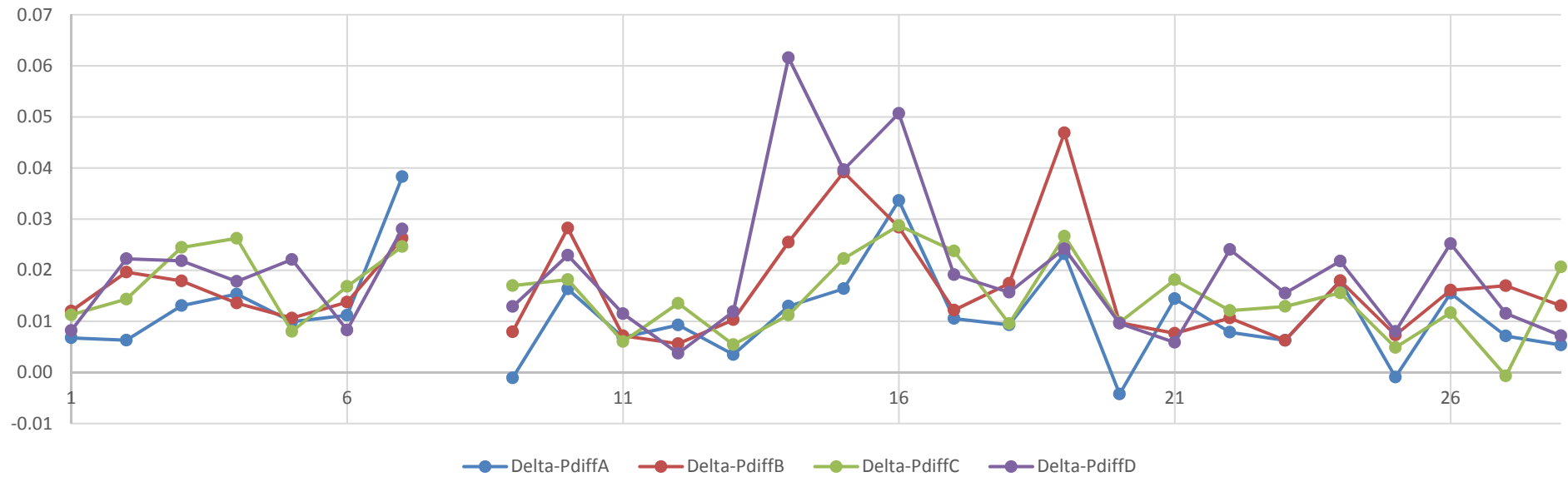


Variations in $\langle G_{\text{mod}} \rangle$ & Frequency Dependence

Pdiff Ratio (Moon/Sky)

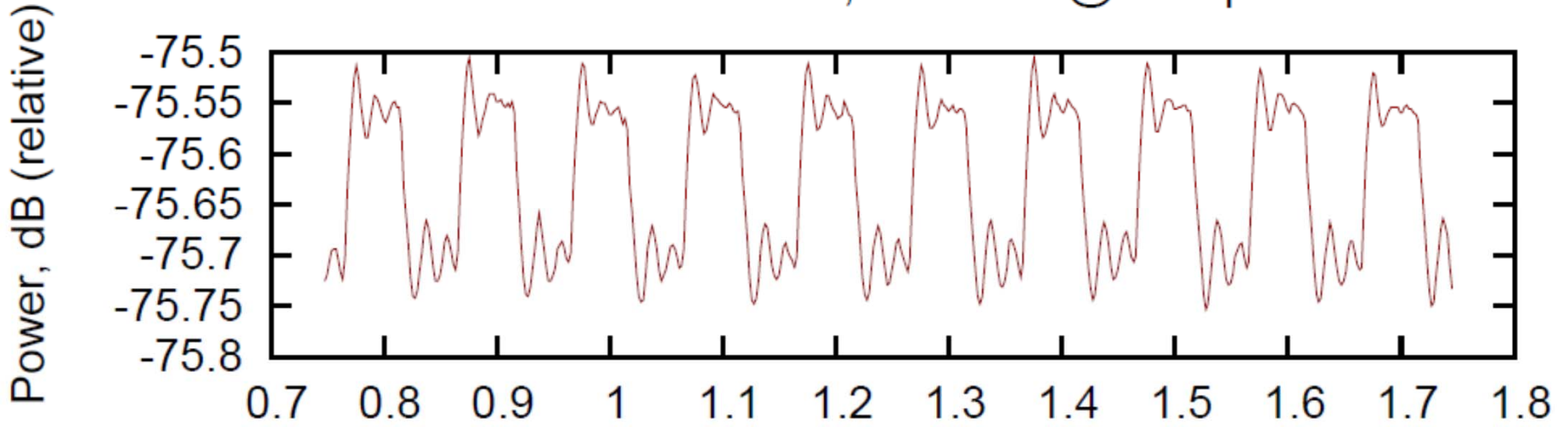


Delta Pdiff (counts)

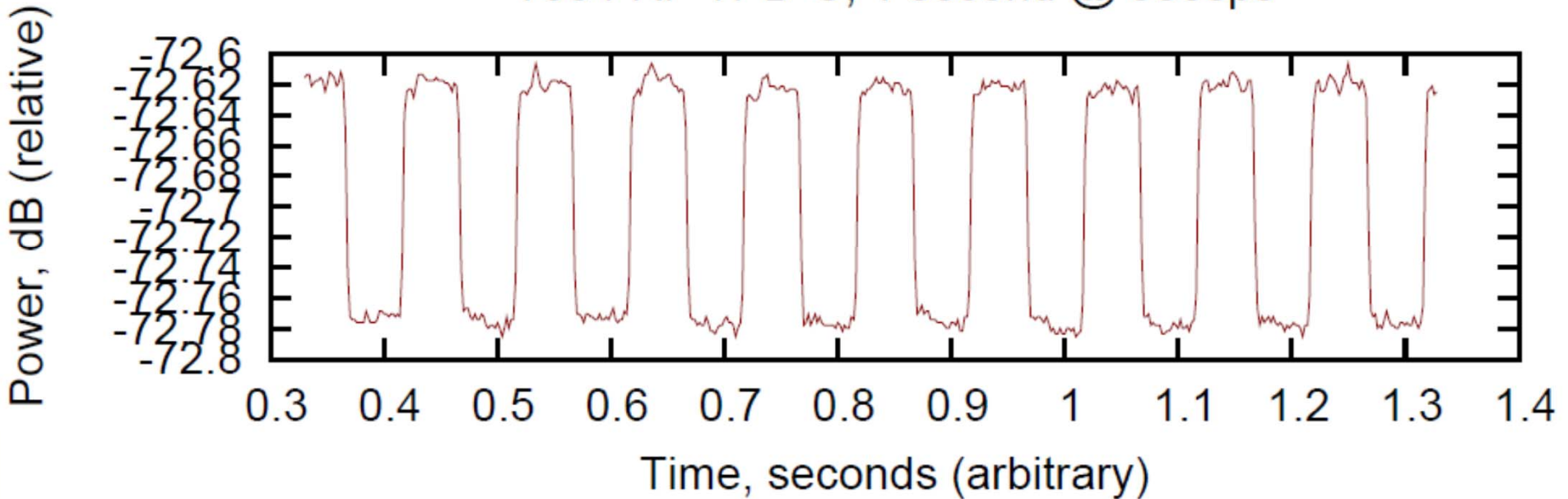


SNI4 S-Band – Bias Card Capacitors

T304 RF TPD-C, 1 second @ 500sps



T304 RF TPD-C, 1 second @ 500sps



HEMT Bias & Frequency Dependence

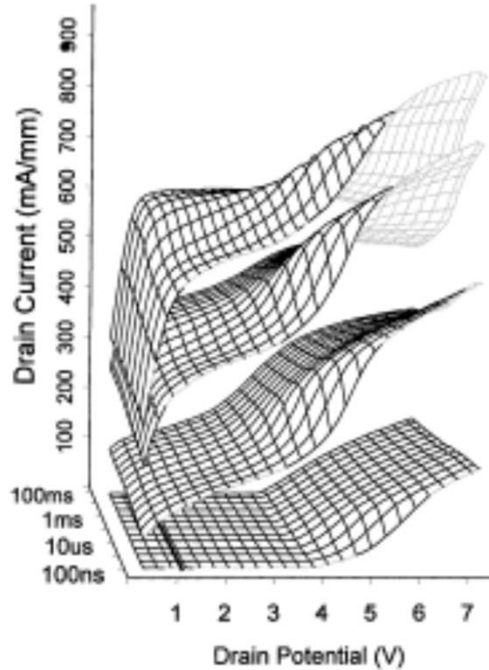


Fig. 3. Time-evolution view of the characteristics shown in Fig. 2.

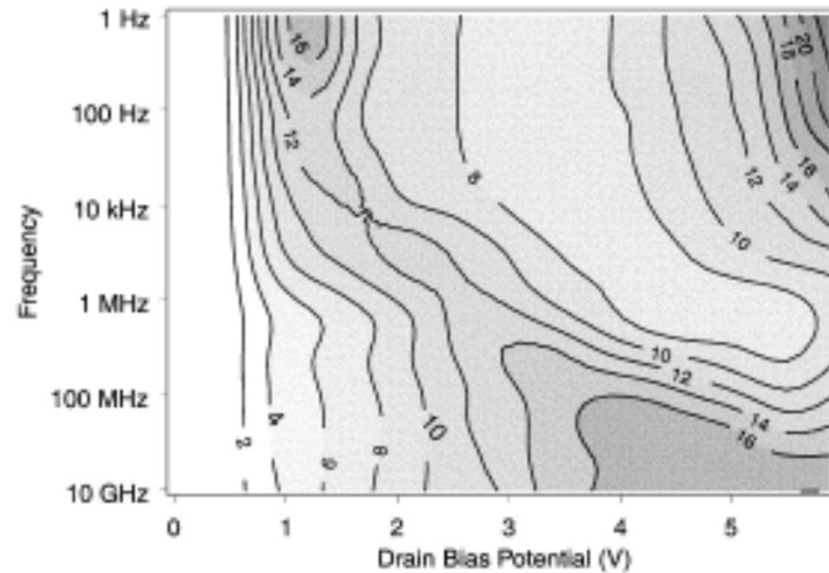


Fig. 4. Intrinsic gain versus drain bias for the HEMT in Fig. 1 operating at 25 °C with $V_{GS} = -0.4$ V. Darker areas correspond to regions of higher gain.

Parker & Rathmell, Bias and Frequency Dependence of FET Characteristics.

IEEE Transactions on Microwave Theory and Techniques, Vol 51, No. 2, Feb 2003

Conclusions & Next Steps

- Main source of the gain compression is likely in the antenna analog electronics, before the attenuator at the input to the T304.
- Majority of the P_{diff} compression at X-band may be attributable to the gain modulation, which is likely a modulation of the LNA or post-amp bias.
- Differences *within* A-B and C-D pairs are partly due to differences in cal power as a function of frequency, but gain modulation appears to be frequency dependent too.
- Focus investigations into gain modulation as a function of antenna power. (Modulation gets progressively worse with increased antenna power.)
- Follow up test with T304s tuned to same frequency (isolate T304 tuning from any frequency-dependent effects in the front end). Include multiple spectral windows within each IF.



Next Steps

- Work with Front End Group to inspect an Rx - Hot load test.
 - Attempt to see gain modulation at a selected X-band front-end output and intermediate gain stages.
 - Look for modulation in bias signals to the LNAs and post-amps.
 - Look for changes in gnd. and other reference voltages
 - Ground loops?
 - Remove filter caps on bias lines and note changes.
- Work with Paul Demorest to look at the gain modulation with the correlator time domain tools (e.g., pulsar binning mode).
 - Quantify magnitude of Cal ringing / gain modulation.
 - Confirm that magnitude of the modulation is affected by Pantenna (vs. just Pcal radiated by other receivers).
 - Look for correlation between gain changes and Pdiff compression.
 -



www.nrao.edu
science.nrao.edu
public.nrao.edu

*The National Radio Astronomy Observatory is a facility of the National Science Foundation
operated under cooperative agreement by Associated Universities, Inc.*

May 19, 2016

ngVLA Overview

